

**1 The energetics of melting fertile heterogeneities**  
**2 within the depleted mantle**

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3 **Abstract.** To explore the consequences of mantle heterogeneity for pri-  
4 mary melt production, we develop a mathematical model of energy conser-  
5 vation for an upwelling, melting body of recycled oceanic crust embedded  
6 in the depleted upper mantle. We consider the end-member geometric cases  
7 of spherical blobs and tabular veins. The model predicts that thermal dif-  
8 fusion into the heterogeneity can cause a factor-of-two increase in the degree  
9 of melting for bodies with minimum dimension smaller than  $\sim 1$  km, yield-  
10 ing melt fractions between 50 and 80%. The role of diffusion is quantified  
11 by an appropriately defined Peclet number, which represents the balance of  
12 diffusion-driven and adiabatic melting. At intermediate Peclet number, we  
13 show that melting a heterogeneity can cool the ambient mantle by up to  $\sim 20$   
14 K (spherical) or  $\sim 60$  K (tabular) within a distance of two times the char-  
15 acteristic size of the body. At small Peclet number, where heterogeneities are  
16 expected to be in thermal equilibrium with the ambient mantle, we calcu-  
17 late the energetic effect of pyroxenite melting on the surrounding peridotite;  
18 we find that each 5% of recycled oceanic crust diminishes the peridotite de-  
19 gree of melting by 1–2%. Injection of the magma from highly molten bod-  
20 ies of recycled oceanic crust into a melting region of depleted upper man-  
21 tle may nucleate reactive-dissolution channels that remain chemically iso-  
22 lated from the surrounding peridotite.

## 1. Introduction

23 It is well established that the mantle is chemically heterogeneous on length-scales smaller  
24 than the 100 km typical size of the melting region beneath a plate-boundary or hot-spot  
25 volcano [e.g. Hofmann, 1997]. Despite this, most studies of melting at plate tectonic  
26 boundaries and hot spots still treat the mantle as a homogeneous source. Such models,  
27 irrespective of their sophistication, probably miss crucial aspects of basalt petrogenesis;  
28 they may therefore lead to incorrect inversions of geochemical data for mantle properties  
29 and processes. Treatment of melting of a lithologically heterogeneous mantle remains  
30 largely qualitative, lacking the rigor necessary for quantitative geochemical modeling.  
31 The present work seeks to address this deficiency.

32 A range of studies have inferred that heterogeneities are formed from previously sub-  
33 ducted oceanic and continental crust and lithosphere [e.g. Hofmann and White, 1982;  
34 Willbold and Stracke, 2010] that has been stirred into the mantle over geologic time.  
35 This stirring continuously reduces the characteristic size of heterogeneities [Hoffman and  
36 McKenzie, 1985; Allègre and Turcotte, 1986], and increases the rate of homogenization  
37 by solid-state diffusion. The surviving heterogeneities are those that remain larger than  
38 the decimeter scale over which diffusion would erase chemical variations on the billion-  
39 year time-scale of mantle convection [Hofmann and Hart, 1978]. Many of these have a  
40 pyroxene-rich lithology that may contain garnet, spinel, and other accessory phases, but  
41 has little or no olivine. In this manuscript, we do not attempt to distinguish between the  
42 pyroxenites, garnet pyroxenites, and eclogites that are included in this range of lithologies;  
43 rather we generalize them to a class of mantle heterogeneity composed of fertile, recycled

44 crustal rocks. We do not directly address other classes of mantle heterogeneity such as  
45 recycled sediments.

46 Within the class considered here, the size-distribution and concentration of hetero-  
47 geneities throughout the mantle remains the subject of controversy. Geochemical mea-  
48 surements and models put constraints on the proportion of enriched materials in the  
49 source regions of mid-ocean ridges. Hirschmann and Stolper [1996] used a variety of  
50 trace-element and isotopic measurements to estimate that the MORB-source contains 3–  
51 6% garnet pyroxenite. Melting experiments by Pertermann and Hirschmann [2003a] more  
52 accurately determined the productivity of upwelling pyroxenite, and lead them to suggest  
53 2–3% pyroxenite, though they note that this result is sensitive to the assumed fertility  
54 of the pyroxenite. Sobolev et al. [2007] studied trace element and forsterite content in  
55 olivine phenocrysts from MORB, OIB, komatiites, and within-plate lavas and suggested  
56 that variations in Mn and Ni content, among other indicators, call for about 5% recycled  
57 oceanic crust in the MORB source and 20% in the OIB source. This is approximately  
58 consistent with estimates by Ito and Mahoney [2005], determined by modeling the melting  
59 process for ridges and plumes, and seeking the range of starting compositions that can  
60 explain both OIB and MORB geochemical and isotopic systematics. Each of these studies  
61 used a different approach to quantify the contribution of recycled oceanic crust, yet their  
62 results are similar; it seems reasonable to expect that enriched heterogeneities compose  
63 less than one fifth of the mantle that melts at most ridges and hot spots.

64 Constraining the sizes and shapes of mantle heterogeneities has proven more difficult  
65 than determining their relative proportion. Seismic scattering techniques described by  
66 Helffrich [2006] provide support for blobs of mineralogically distinct heterogeneities with

67 a characteristic scale of  $\lesssim 10$  km, distributed uniformly throughout the mantle. Kogiso  
68 et al. [2004] argued for a size of  $\gtrsim 1$  m, which they calculated based on the observational  
69 constraint of radiogenic osmium signatures in rocks recovered from mid-ocean ridges and  
70 hot-spots. Yasuda and Fujii [1998] noted that the negative buoyancy of eclogite blobs  
71 means that blobs of diameter 40 km or larger cannot ascend through the upper mantle,  
72 though the density that they assumed for eclogite may have been too large [Pertermann  
73 and Hirschmann, 2003b]. Others have argued for heterogeneity of equal magnitude at all  
74 scales based on observed geographical variation in the isotopic ratios of lavas [e.g. Gurnis,  
75 1986].

76 Since fertile heterogeneities are a small fraction of mantle volume, and are apparently  
77 dispersed at sizes smaller than 10 km, we study their melting systematics using an idealized  
78 model of an isolated heterogeneity. Such a heterogeneity would initially be in thermal  
79 equilibrium with the surrounding mantle, because the rate of thermal diffusion is fast  
80 relative to that of convective overturning. Because of its composition, it would begin  
81 decompression melting at higher pressures than the ambient, more depleted mantle. This  
82 melting requires energy to convert solid to liquid, and it diminishes the temperature of the  
83 blob relative to the surrounding mantle. The associated temperature gradient drives heat  
84 flow into the blob, and increases the melting rate, while cooling the surrounding mantle.  
85 What is the magnitude of this effect? This question was first addressed by Sleep [1984],  
86 who developed models of heat conduction into a fertile, melting sphere and tabular body.

87 Sleep [1984] considered a fertile heterogeneity undergoing decompression melting at an  
88 upwelling rate of 3 cm/yr and calculated the size it must have such that the characteristic  
89 thermal diffusion time is approximately equal to its duration in the melting column,

90 obtaining a value of 5 km. Moreover, he recognized that heterogeneities much smaller than  
91 this size would melt in thermal equilibrium with the surrounding mantle, while those much  
92 larger would melt in thermal isolation. Sleep [1984] then derived a mathematical model  
93 of partial melting of such heterogeneities based on diffusive heat flow and parameterized  
94 thermodynamic properties. Analysis of this model lead him to conclude that thermal  
95 diffusion could cause an enhancement of melting by up to a factor of seven over the  
96 thermally isolated case. This extremely large enhancement contrasts with the results that  
97 we present below.

98 In the present manuscript we take a similar approach to Sleep [1984]—we consider ther-  
99 mal diffusion into a fertile heterogeneity of spherical or tabular shape—but we provide a  
100 more comprehensive study and obtain significantly different results. Our approach is more  
101 rigorous in that we derive an equation for conservation of energy that explicitly couples  
102 melting with thermal diffusion. In contrast to Sleep [1984], we obtain full analytical solu-  
103 tions in terms of nondimensional parameters, and use them to elucidate the controls on  
104 and limits of partial melting of heterogeneities. Furthermore, on empirical [Pertermann  
105 and Hirschmann, 2003a] and theoretical grounds [Hirschmann et al., 1999], we consider  
106 the nonlinear relationship between melt fraction and temperature, where Sleep [1984]  
107 assumed linearity. For the nonlinear case, we rely on numerical solutions of the govern-  
108 ing equations; these are documented and validated in Appendix B. Our calculations are  
109 based on parameter values from published experiments on G2 pyroxenite [Pertermann and  
110 Hirschmann, 2003a], which has a composition thought to approximate recycled oceanic  
111 crust. We compare the results of these calculations with those obtained by Sleep [1984],  
112 who used different and probably unrealistic parameters for pyroxenite melting (for exam-

113 ple, Sleep [1984] assumed a near-solidus isobaric productivity  $\partial F/\partial T|_P$  that is a factor  
114 of five to twenty times larger than the value obtained empirically by Pertermann and  
115 Hirschmann [2003a]).

116 Our work is also related to that of Phipps Morgan [2001], who considered narrow,  
117 multi-layered tabular veins of fertile pyroxenite within depleted peridotite, and computed  
118 melting behavior for a variety of compositional scenarios. He assumed thermal equilibrium  
119 between all the layers in the model. Similar models were developed by Hirschmann and  
120 Stolper [1996] and Stolper and Asimow [2007]. And whereas these authors' work focused  
121 on the effects of melting a compositional *mélange* with variable solidus (and solidus slope)  
122 between solid phases, here we are interested in accounting for the finite time-scale of  
123 thermal equilibration. Our model is not restricted to any one set of pyroxenite melting  
124 parameters, though in this manuscript we limit consideration to a single example that we  
125 hope is representative.

126 The relative simplicity of the physical model is based on a number of key assumptions.  
127 The most important of these is that heterogeneities are isolated from one another, such  
128 that their thermal disturbances do not interact. We further assume that the characteristic  
129 size of a heterogeneity is small relative to the vertical distance between the depth at  
130 which it reaches its solidus and the depth of the ambient mantle solidus. This latter  
131 assumption allows us to approximate the lithostatic pressure as being constant within the  
132 heterogeneity. We assume that heterogeneities upwell at the same speed as the ambient  
133 mantle; other authors have considered the possibility that chemically dense, unmolten,  
134 recycled material upwells at a slower speed due to its negative buoyancy [Yasuda and  
135 Fujii, 1998; Pertermann and Hirschmann, 2003b]. Finally, we assume that melt does

136 not segregate from the host rock, which is the thermodynamic equivalent of assuming  
137 batch melting. If the ambient mantle remains unmolten as an embedded heterogeneity  
138 partially melts, magma may be held within the heterogeneity by a permeability barrier.  
139 Alternatively, it may chemically react with the surrounding rock [Yaxley and Green, 1998;  
140 Kogiso et al., 2004], propagate through it by diking, or open the pores through surface-  
141 energy driven flow [Riley and Kohlstedt, 1991]. Even a leaky permeability barrier around  
142 a fertile enclave would significantly restrict magmatic segregation if the scale of the sphere  
143 is smaller than the compaction length [Spiegelman, 1993]. For the case of a tabular vein,  
144 these arguments are clearly tenuous; we therefore justify the assumption on grounds of  
145 mathematical convenience, and flag this as an issue to be remedied in future work. These  
146 assumptions enable us to formulate a mathematical model that admits analytical solution  
147 in some cases, and high-accuracy, efficient numerical solutions in others. We can thus  
148 provide a thorough exploration of the model behavior, and draw conclusions that will  
149 inform more detailed work in which assumptions are relaxed.

150 Although this manuscript focuses on the energetics of pyroxenite melting, the results  
151 also have geochemical implications. It is well-known that garnet is an important phase  
152 in recycled oceanic crust, and that it has a unique set of affinities for trace elements [e.g.  
153 Stracke et al., 1999; Pertermann et al., 2004]. The trace-element budget of melts derived  
154 from pyroxenite is therefore sensitive to the presence of residual garnet; the persistence of  
155 garnet in the residue depends on pressure, but it also depends on the degree of melting.  
156 The work presented here provides a framework for evaluating the degree of melting of  
157 pyroxenite enclaves within the depleted, peridotite upper mantle; it also highlights the  
158 difficulties in inverting geochemical data for the physical characteristics of mantle hetero-



159 geneity by showing that the size and shape of heterogeneities has a significant influence on  
160 their melting behavior. Perhaps a more significant difficulty is that the pathways of melt  
161 transport from deep-melting, fertile heterogeneities to the surface are poorly constrained,  
162 and it is increasingly understood that the details of melt transport have a crucial impact  
163 on observed geochemical patterns [e.g. Spiegelman and Kelemen, 2003; Liang et al., 2011].

164 The next section describes the idealized physical scenario to be analyzed, and lays  
165 out a mathematical formulation of the problem. The Results section then presents non-  
166 dimensional output from analytical and numerical solutions. The Discussion section ex-  
167 amines these results in more detail and provides a subset of them in dimensional form. We  
168 summarize our findings briefly at the end, and give the details of analytical and numerical  
169 methods in two appendices. Throughout this manuscript, the focus is on the slightly  
170 more complicated case of a spherical heterogeneity, however in the Discussion section, we  
171 compare our results for spherical and tabular bodies.

## 2. The mathematical model

172 Figure 1 is a schematic representation of the model set-up. We consider two end-member  
173 cases: a spherical blob of radius  $R$  and an infinite tabular vein of half-width  $R$ . Both are  
174 made of a uniform, fertile lithology, and are embedded within the depleted upper mantle,  
175 upwelling at speed  $W > 0$ . We develop the mathematical model using the case of the  
176 spherical blob and present the case of a tabular vein in Section 2.3.

177 We assume that the blob is small enough that we can neglect vertical variation of all  
178 properties within it. This assumption is awkward for the tabular vein, but not unreason-  
179 able if we consider that temperature gradients (and hence heat flow) normal to the vein  
180 will greatly exceed those within it. As the heterogeneity ascends, it reaches a pressure

181  $p_0 = \rho g z_0$  where its temperature is equal to the solidus temperature  $T_0$ ; we label this  
 182 moment  $t = 0$  ( $\rho$  is the density of all materials under consideration;  $g$  is the acceleration  
 183 due to gravity; both are assumed constant). The ambient mantle is also upwelling with  
 184 speed  $W$ , and it reaches its solidus temperature  $T_1$  at a shallower pressure  $p_1 = \rho g z_1$  (and  
 185 hence at a later time  $t = t_1$ ). We investigate the time interval between 0 and  $t_1$ , when the  
 186 blob is partially molten but the ambient mantle is entirely solid.

Motivated by the experimental results of Pertermann and Hirschmann [2003a], the solidus and liquidus temperatures of the blob are taken to depend on pressure only:

$$T_s(p) = T_0 + \gamma^{-1}(p - p_0), \quad (1a)$$

$$T_l(p) = T_0 + \gamma^{-1}(p - p_0) + \Delta T, \quad (1b)$$

where  $\Delta T$  is a constant and  $\gamma$  is the Clapeyron slope. Because of the uniform upward motion, the pressure experienced by the blob changes with time according to

$$p(t) = p_0 - \rho g W t. \quad (2)$$

187 The coordinate system is fixed to the center of the spherical heterogeneity, which has  
 188 a radius  $R$ . The goal is to determine the temperature field  $T(r, t)$  outside the blob as  
 189 it melts, and to calculate the average temperature and melting rate within the blob.  
 190 Magmatic segregation can be neglected if the blob is much smaller than the compaction  
 191 length within it. This will be true for smaller values of  $R$ , but since we are concerned  
 192 only with the averaged melting properties of the blob, we expect that mass redistribution  
 193 within the blob would have only a small effect on our results.

With the above assumptions, conservation of energy is

$$\frac{\partial H}{\partial t} + \rho g W = k \nabla^2 T, \quad (3)$$

where  $\rho g W$  is the rate of change of potential energy and

$$dH = \rho L dF + \rho c_p dT + (1 - \alpha T) dp \quad (4)$$

represents an infinitesimal change in bulk enthalpy in terms of its contributing parts.  $F$

here is the volume fraction of melt, and the remaining symbols are defined in Table 1.

Combining (2), (3), and (4) gives

$$\rho L \frac{\partial F}{\partial t} + \rho c_p \frac{\partial T}{\partial t} - \alpha T \frac{\partial p}{\partial t} = k \nabla^2 T. \quad (5)$$

We will non-dimensionalise with the following scales

$$[x] = R, \quad [t] = \frac{\Delta p}{\rho g W}, \quad (6)$$

and define the non-dimensional temperature and pressure as

$$\theta = \frac{\gamma}{\Delta p} (T - T_0), \quad P = \frac{p - p_0}{\Delta p}, \quad (7)$$

where  $\Delta p = p_0 - p_1$  and the temperature scale is the change in the solidus temperature from  $p_0$  to  $p_1$ . Using these scales and linearizing the adiabatic gradient about  $T = T_0$  gives

$$\mathcal{S} \frac{\partial F}{\partial t} + \frac{\partial \theta}{\partial t} + \mathcal{A} = \frac{1}{\text{Pe}} \nabla^2 \theta, \quad (8)$$

where the Peclet number is

$$\text{Pe} = \frac{\tau_R}{\tau_a} = \frac{\rho g W R^2}{\kappa \Delta p},$$

the Stefan number is

$$\mathcal{S} = \frac{L \gamma}{c_p \Delta p},$$

and the adiabatic parameter is

$$\mathcal{A} = \frac{\alpha \gamma T_0}{\rho c_p}.$$

194 All symbols in equation (8) represent non-dimensional quantities. The Peclet number  
 195 is the ratio of the time-scale  $\tau_R = R^2/\kappa$  for diffusion of heat across the blob to the  
 196 time-scale  $\tau_a = \Delta p/(\rho g W)$  for advection of the blob from  $p_0$  to  $p_1$ . It is the principal  
 197 control parameter in the problem; for  $\text{Pe} \rightarrow 0$  the diffusive heat transport dominates  
 198 the thermal budget and the blob is in thermal equilibrium with the surrounding mantle,  
 199 while for  $\text{Pe} \rightarrow \infty$ , diffusion is negligible relative to advective transport, and the blob  
 200 melts adiabatically.

201 The adiabatic parameter is the linearized rate of energy consumption by adiabatic  
 202 expansion. We neglect the term associated with the adiabatic parameter in what follows.  
 203 In A5 we show that incorporating this term gives rise to the usual adiabatic temperature  
 204 gradient, and introduces a factor of  $(1 - \mathcal{A})$  into the melt productivity.

With the above scaling, the equations for the solidus and liquidus (1) become

$$\theta_s = P = -t, \quad (9a)$$

$$\theta_l = P + \Delta\theta = -t + \Delta\theta, \quad (9b)$$

205 where  $\Delta\theta = \gamma\Delta T/\Delta p$  is the non-dimensional, constant temperature offset between solidus  
 206 and liquidus. The non-dimensional pressure is given by  $P(t) = -t$  with  $0 \leq t \leq 1$ .

207 We can put rough constraints on all of the material parameters, as given in Table 1, but  
 208 cannot prescribe the radius of the blob or its upwelling rate. However, these latter two  
 209 parameters are combined in the Peclet number, so we need explore the variation in only a  
 210 single dimensionless parameter. Note that a value of  $\Delta\theta$  larger than unity indicates that  
 211 the temperature difference between the solidus and the liquidus of the blob is larger than  
 212 the temperature difference between its solidus for non-dimensional pressures  $P = 0$  and  
 213  $P = -1$ . This means that for  $\Delta\theta > 1$  and  $t \leq 1$ , the blob *cannot* melt to  $F = 1$ . In the

214 current work we do not consider the case of a fully molten heterogeneity, which can only  
 215 occur for  $t \leq 1$  if  $\Delta\theta \leq 1$ .

## 2.1. Within the blob

216 Since we are interested in melting of a fertile blob of recycled material, we first formulate  
 217 equations that capture its melting properties.

### 2.1.1. Simplified melting relations for fertile, recycled oceanic crust

When the temperature within the blob is above the solidus, a dimensionless homologous temperature is

$$\Theta = \frac{\theta - \theta_s}{\Delta\theta}, \quad \text{when } \theta_s \leq \theta \leq \theta_l. \quad (10)$$

Pertermann and Hirschmann [2003a] found, for the anhydrous pyroxenite composition G2 (N.B. Pertermann and Hirschmann [2003a] use the term pyroxenite to describe all pyroxene-rich, olivine-poor mantle heterogeneities), that the relationship between the non-dimensional homologous temperature  $\Theta$  and the degree of melting can be represented as

$$F = a\Theta + (1 - a)\Theta^2, \quad (11)$$

219 where  $0 \leq a \leq 1$ . Their data was best fit for  $a \approx 1/4$ . Their pyroxenite composition,  
 220 chosen to be similar to typical oceanic crust, cannot represent *all* flavors of mantle het-  
 221 erogeneity. It is, instead, a well-characterized and important example. We assert that the  
 222 melting consequences of other compositions can be investigated within the mathematical  
 223 framework established below, by repeating our calculations with modified values for ma-  
 224 terial constants. For example, one could model the presence of volatile elements in the  
 225 pyroxenite by modifying parameters  $T_0$ ,  $p_0$  and  $a$  to produce a “tail” of low- $F$  melting at  
 226 high pressure [Hirschmann et al., 1999].

227 **2.1.2. The degree of melting of the blob**

To simplify the analysis, we now assume a homogeneous distribution of temperature and melt fraction within the blob, and define these as  $\theta_B$  and  $F_B$  respectively. This approximation will hold at small Peclet numbers; in Section 3.3 we relax this constraint and consider radially variable blobs. Proceeding with the averaged quantities  $\theta_B$  and  $F_B$ , we can integrate equation (8) over the non-dimensional volume of the blob  $V = 4\pi/3$ . This gives

$$\mathcal{S} \frac{\partial F_B}{\partial t} + \frac{\partial \theta_B}{\partial t} = \frac{3}{4\pi} \frac{1}{\text{Pe}} \int_S \nabla \theta \cdot \mathbf{dS}, \quad (12)$$

where we have used the divergence theorem to convert the volume integral into a surface integral. We can evaluate this integral in spherical coordinates as the spherically symmetric gradient in the radial direction times the non-dimensional surface area of the blob  $4\pi$ . We then integrate with respect to time to obtain

$$\mathcal{S} F_B + \theta_B = \frac{3}{\text{Pe}} \int_0^t \left. \frac{\partial \theta}{\partial r} \right|_{1,\tau} d\tau. \quad (13)$$

Both  $\theta_B$  and  $F_B$  can be expressed in terms of  $\Theta_B$  by using (10) and (11). Equation (13) becomes

$$(\Delta\theta + a\mathcal{S})\Theta_B + (1-a)\mathcal{S}\Theta_B^2 = t + \frac{3}{\text{Pe}} \int_0^t \left. \frac{\partial \theta}{\partial r} \right|_{1,\tau} d\tau, \quad (14)$$

which relates the homologous temperature inside the blob to the integrated heat flux into the blob from the ambient mantle. The above equation can be solved for  $\Theta_B$ ; the result is

$$\Theta_B = \frac{\sqrt{1 + 4\Lambda f_B} - 1}{2\Lambda}, \quad (15)$$

where

$$\Lambda = \frac{(1-a)\mathcal{S}}{\Delta\theta + a\mathcal{S}}, \quad f_B = \frac{1}{a\mathcal{S} + \Delta\theta} \left( t + \frac{3}{\text{Pe}} \int_0^t \left. \frac{\partial \theta}{\partial r} \right|_{1,\tau} d\tau \right). \quad (16)$$

In the linear limit where  $a \rightarrow 1$  ( $\Lambda \rightarrow 0$ ), equations (15) and (16) simplify to

$$\Theta_B = f_B = \frac{1}{\mathcal{S} + \Delta\theta} \left( t + \frac{3}{\text{Pe}} \int_0^t \frac{\partial\theta}{\partial r} \Big|_{1,\tau} d\tau \right). \quad (17)$$

228 This linear case can be treated analytically; details are provided in A.

## 2.2. Outside the blob

Outside the blob, the mantle is below its solidus and hence  $F = 0$ . Equation (8) becomes

$$\frac{\partial\theta}{\partial t} = \frac{1}{\text{Pe}} \frac{1}{r^2} \frac{\partial}{\partial r} \left( r^2 \frac{\partial\theta}{\partial r} \right), \quad (18)$$

229 where we have chosen spherical coordinates and used the symmetry of the problem to  
230 discard terms.

Equation (18) has boundary conditions

$$\theta(1, t) = -t + \Theta_B \Delta\theta, \quad (19a)$$

$$\theta(\infty, t) = 0, \quad (19b)$$

where  $\Theta_B$  is given by (15) and (16). The first of these boundary conditions represents the continuity of temperature at the surface of the blob, and the second represents the constant far-field temperature of the ambient mantle. The initial condition is uniform temperature,

$$\theta(r, 0) = 0. \quad (20)$$

## 2.3. A tabular heterogeneity

The governing equations are simplified slightly if we consider a tabular heterogeneity of half-width  $R$  and infinite extent; equations (14) and (18) are replaced by

$$(\Delta\theta + a\mathcal{S})\Theta_B + (1 - a)\mathcal{S}\Theta_B^2 = t + \frac{1}{\text{Pe}} \int_0^t \left. \frac{\partial\theta}{\partial x} \right|_{1,\tau} d\tau, \quad (21)$$

$$\frac{\partial\theta}{\partial t} = \frac{1}{\text{Pe}} \frac{\partial^2\theta}{\partial x^2}, \quad (22)$$

where we have taken the  $x$ -axis in the direction normal to the tabular body. The boundary conditions (19) are unchanged, except for use of the modified solution for  $\Theta_B$ . As with the spherical blob, we neglect magmatic segregation for the tabular body. A model similar to this was considered by Sleep [1984]. In the Results section, we limit our attention to solutions for the spherical blob. The tabular vein is reintroduced in the Discussion section, where we examine the consequences of blob shape on melting behavior, for these two end-member cases.

For a spherical (or tabular) blob, the solution is obtained by solving the system of equations (14) and (18) (or (21) and (22)), with (16), (19), and (20) on the domain  $1 \leq (r, x) < \infty$  and  $0 \leq t \leq 1$  for given values of  $a$ ,  $\text{Pe}$ ,  $\mathcal{S}$ , and  $\Delta\theta$ .

In the end-member cases of melting for infinite and zero Peclet number corresponding, respectively, to purely adiabatic and purely isothermal melting, it is not necessary to solve for the temperature structure outside the blob [Sleep, 1984]. Figure 2 shows a schematic representation of the temperature–pressure path taken by the blob in these two cases, and illustrates the thermodynamic phase diagram (for  $a = 1$ ). Solutions for intermediate Peclet numbers follow paths between the end-member curves, but must be obtained through analytical or numerical solutions to the governing equations. These are presented in the next section.



### 3. Results

249 In this section we present results from both analytical and numerical solutions of the  
 250 governing equations for a spherical blob. We fix all the dimensionless parameters except  
 251 the Peclet number, which captures the variation in both upwelling rate and blob size.  
 252 Solutions are presented for a range of Peclet numbers.

#### 3.1. Linear melting approximation

253 When  $a = 1$ , the boundary condition described by (17) and (19a) is linear and an  
 254 analytical solution can be obtained using Laplace transforms (see A1). These are plotted  
 255 in Figure 3 for different values of the Peclet number. Panel (a) shows the degree of  
 256 melting of the blob as a function of time. The range of curves is bounded by dashed  
 257 lines for solutions at asymptotic values of  $Pe$ . When  $Pe \rightarrow 0$ , the blob is in thermal  
 258 equilibrium with its surroundings. In this case, the degree of melting is determined entirely  
 259 by the Clapeyron slope and the relationship between temperature and degree of melting  
 260 in (11). Blobs of radius  $\mathcal{O}(1 \text{ km})$  or smaller would behave according to the small-Peclet  
 261 limit. When  $Pe \rightarrow \infty$ , the blob is a closed thermodynamic system, exchanging no heat  
 262 with the surrounding mantle. In this case, the degree of melting is controlled by the  
 263 decreasing solidus temperature and the latent-heat cost of melting. Blobs of radius  $\mathcal{O}(100$   
 264  $\text{ km})$  or larger would behave according to the large-Peclet limit. Further dimensional  
 265 considerations are deferred to the discussion section, below.

For finite values of  $Pe$ , the ambient thermal state,  $\theta(r, 0) = 0$ , is altered by diffusion  
 of heat into the relatively cool blob. The magnitude of the disturbance is plotted in  
 Figure 3b for  $t = 1$ . The amplitude of the near-field disturbance grows with increasing  
 $Pe$ , while the decay length of the disturbance decreases. In other words, when diffusion is

unimportant ( $\text{Pe} \rightarrow \infty$ ), temperature differences from the background are large, but the radius of the thermal halo is small. When diffusion is important ( $\text{Pe} \rightarrow 0$ ), the opposite is true. In the latter case, where diffusive equilibrium is reached in infinitesimal time, we can make the quasi-steady approximation, setting  $\partial\theta/\partial t \approx 0$  in (18) to obtain

$$0 \approx \frac{1}{r^2} \frac{\partial}{\partial r} \left( r^2 \frac{\partial \theta}{\partial r} \right). \quad (23)$$

266 Inspection of this equation shows that solutions must take the form  $\theta(r, t) = \theta_B(t)/r$ . This  
 267 predicts that at small Peclet number, the thermal disturbance close to the blob should  
 268 scale as  $r^{-1}$  (Figure 3).

### 3.2. Nonlinear melting

Melting experiments on pyroxenite by Pertermann and Hirschmann [2003a] are best fit with  $a \approx 1/4$ , so we now examine solutions for that case. Since equation (14) is then nonlinear, we must rely on numerical methods to solve the problem (see B). Figure 4 shows the results of a suite of calculations for different values of  $\text{Pe}$ , with  $a = 1/4$ . As with the case of  $a = 1$ , the curves for  $F_B(t)$  are bounded above and below by asymptotes for  $\text{Pe} \rightarrow 0$  and  $\text{Pe} \rightarrow \infty$ , respectively. The latter case, representing adiabatic melting, results in a temperature evolution given by

$$\Theta_B(t; \text{Pe} \rightarrow \infty) = \frac{1}{2\Lambda} \left( \sqrt{1 + \frac{4\Lambda t}{\Delta\theta + a\mathcal{S}}} - 1 \right), \quad (24)$$

which can be obtained from (15) and (16) by neglecting the term describing conduction of heat into the blob. The  $\text{Pe} \rightarrow 0$  case represents melting in thermal equilibrium with the surrounding mantle, and hence at the fixed temperature  $\theta = 0$ . Using (9a) and the definition of  $\Theta$  in (10) we obtain

$$\Theta_B(t; \text{Pe} \rightarrow 0) = \frac{t}{\Delta\theta}. \quad (25)$$

Equations (24) and (25) can be substituted into the melting relationship (11); the results are plotted in Figure 4 as a red, dashed lines.

Figure 4b shows that the perturbation to the ambient temperature field surrounding the blob is not significantly different from the linear case. For small Peclet numbers, the temperature perturbation falls off as  $r^{-1}$ , in accordance with a quasi-steady approximation (23) of the diffusion equation.

### 3.3. Radially variable melting within the blob

At small Peclet number, diffusion efficiently transports heat to the centre of the blob, and neutralizes any temperature gradients. In contrast, when the Peclet number is  $\mathcal{O}(1)$ , the time-scale for diffusion of heat across the blob is comparable to the time for vertical advection through the domain. In this case, we expect a significant temperature gradient across the radius of the blob, with warmer temperatures at the edge and cooler temperatures at the centre. For larger Peclet numbers, diffusive heat transport is inefficient, and the temperature gradient does not penetrate far into the blob, but rather yields higher temperatures and melt fractions in a rim at the edge of the domain. To capture the behavior for  $Pe \gtrsim 1$ , we relax the assumption of homogeneous properties  $F_B(t)$  and  $\theta_B(t)$  within the blob, and instead calculate radially variable properties  $F(r, t)$  and  $\theta(r, t)$ . This is the approach taken by Sleep [1984], although his calculations used 1-dimensional Cartesian rather than spherical geometry. In this section we consider only the linear melting relation,  $a = 1$ ; based on the similarity of Figures 3 and 4, we infer that comparable behavior would be observed for the nonlinear case. For  $a = 1$ , the Laplace transform method can be used, with a numerical scheme to invert the transform (details and references in A2).

290 Figure 5 shows the results of calculations for a radially variable blob. Panel (a) illus-  
291 trates the discussion in the preceding paragraph about the structure of the blob at Peclet  
292 numbers around or above unity. The degree of melting shows a radial dependence at  
293  $Pe \approx 1$  and this gradient steepens and localizes with increasing  $Pe$ . Panel (b) shows the  
294 evolution of the radially-resolved degree of melting with time, for  $Pe = 1$ . From an early  
295 state that is nearly homogeneous in  $F$ , the diffusive heat flux penetrates progressively  
296 further into the blob with time, as shown in panel (d), and establishes a temperature  
297 gradient that spans the radius of the blob. Panel (c) compares the final temperature dis-  
298 tribution within and outside the blob for a range of Peclet numbers. Section 4.1 includes  
299 a quantitative comparison of the radially resolved model and the radially averaged model,  
300 showing that differences are relatively small, but are maximized for  $1 \lesssim Pe \lesssim 10$ .

301 We obtain a parallel set of results (not shown) for the case of the tabular vein, with  
302 moderate but systematic differences from those above. Part of the discussion below is a  
303 comparison of results for spherical and tabular models.

#### 4. Discussion

304 In this section we make a detailed examination of results for spherical and tabular  
305 heterogeneities. Before doing that, however, we compare the behavior of the radially-  
306 averaged spherical model (Figure 3 and 4) with the radially-resolved spherical model  
307 (Figure 5). Most of this Discussion section compares results as a function of the Peclet  
308 number, but at the end, we interpret the results in terms of dimensional heterogeneity  
309 size and upwelling rate. At that point, we also provide a quantitative comparison with  
310 the results obtained by Sleep [1984].

#### 4.1. Model behavior as a function of Peclet number

311 Section 3.3 showed that at intermediate and large Peclet numbers, the spherical blob  
 312 can have significant radial structure. What error do we make, then, in assuming a homoge-  
 313 neous distribution of melting *a priori*? Figure 6 shows that in terms of mean quantities,  
 314 this difference is rather small (the radially resolved model has been averaged over the  
 315 spherical blob to produce an *a posteriori* mean). In general, the figure shows that the  
 316 assumption of a homogeneous blob leads to a larger melt production and larger temper-  
 317 ature contrast. This is because imposing homogeneity within the blob is equivalent to  
 318 requiring that diffusion is infinitely fast there, leading to perfect redistribution of energy.  
 319 Without this assumption, the rim of the sphere warms more rapidly than the core, and  
 320 insulates it from inward diffusion of heat. The difference is largest at intermediate Peclet  
 321 numbers because there, as shown in Figure 5, the gradient in  $F$  and  $\theta$  spans much or all  
 322 of the spherical blob radius.

323 The maximum difference for the mean degree of melting between the two models is about  
 324 5%, and indicates that a blob at  $Pe \approx 1$  may melt to  $\sim 60\%$  rather than  $\sim 65\%$  (for  $a = 1$ ).  
 325 In the present idealized context, this difference may be considered insignificant. Similar  
 326 models that assume progressive removal of melt from the blob (e.g. fractional melting),  
 327 or that are concerned with the chemistry of individual “packets” of melt produced within  
 328 the blob, might find the assumption of homogeneity to be problematic, especially for large  
 329 Peclet numbers.

330 How does melting vary as a function of Peclet number? Figure 7 provides an answer  
 331 in terms of the mean degree of melting (panel a) and temperature (panel b) inside the  
 332 heterogeneity. Evidently, for Peclet numbers smaller than  $\sim 10^{-1}$ , diffusion of heat into

333 the spherical blob (black lines & symbols) dominates the melting budget and keeps the  
334 temperature of the blob equal to the ambient temperature. For Peclet numbers larger  
335 than  $\sim 10^2$ , diffusion of heat makes a negligible contribution to the overall budget and  
336 blob temperatures approach their adiabatic limit (although Figure 5a shows that the  
337 diffusive heat flux may be important in a narrow rim at the edge of the blob). In the  
338 intermediate-Peclet regime, diffusion makes a significant but not dominant contribution,  
339 and the final degree of melting and blob temperature are sensitive to the value of Pe.

340 The tabular vein shows a qualitatively similar behavior (blue lines and symbols in Fig-  
341 ure 7). At very large Peclet numbers, where diffusion makes an insignificant contribution  
342 to melting, the spherical blob and the tabular vein melt to equal extents; at very small  
343 Peclet numbers, where diffusion makes a substantial contribution to powering melting,  
344 the two also agree. Differences appear between these extremes, with the tabular vein  
345 exhibiting less extensive melting at any intermediate value of Pe. This can be understood  
346 in terms of the symmetry of each blob shape. In the spherical case, a circular patch on  
347 the surface of the heterogeneity draws a diffusive heat flow from a (truncated) cone of  
348 mantle, its volume increasing with the cube of distance from the blob. In the tabular  
349 case, a circular patch on the surface of the heterogeneity draws heat only from a cylinder  
350 of mantle, with a volume proportional to linear distance from the vein.

351 Figure 7 also shows that the nonlinearity of the temperature–melting function (11) has  
352 a simple and uniform effect on the final degree of melting and the final temperature.  
353 Lower values of  $a$  require higher temperature to reach a given  $F_B$ . With other parameters  
354 held constant, decreasing  $a$  gives smaller  $F_B$  as a function of time, which means less  
355 conversion of sensible to latent heat, and therefore a smaller blob-temperature difference

356 from ambient. Since it is this temperature difference that drives thermal diffusion and  
 357 further melting, the diffusive flow of energy into the blob decreases with decreasing  $a$ .

Despite reductions to maximum  $F_B$  from nonlinear effects, geometry of the heterogeneity, and adiabatic decompression (which gives a  $\sim 10\%$  reduction), melting at low Peclet numbers yields degrees of partial melting that are well in excess of 50%. And while less fertile compositions than the ones considered here would generate smaller  $F_B$ , our prediction should hold if the G2 pyroxenite composition of Pertermann and Hirschmann [2003a] is representative of recycled oceanic crust. Furthermore, Figure 7a shows that even for the nonlinear case, we can approximate the maximum increase in degree of melting due to thermal diffusion by  $1/\lambda$  where

$$\frac{1}{\lambda} = \frac{\mathcal{S} + \Delta\theta}{\Delta\theta} = \frac{L + c_p\Delta T}{c_p\Delta T}, \quad (26)$$

358 which is independent of the temperature difference  $\gamma/\Delta p$  between the solidii of recycled  
 359 oceanic crust and depleted upper mantle (recall that  $L$  is latent heat,  $c_p$  is specific heat capacity, and  $\Delta T$  is the liquidus–solidus temperature interval of pyroxenite). This equation  
 360 is equivalent to eqn. (11) of Sleep [1984].  
 361

362 How broad is the effect of melting within the heterogeneity on the mantle temperature  
 363 around it? Figure 8 addresses this question for a range of Peclet numbers for both the  
 364 spherical blob (black lines and symbols) and the tabular vein (blue lines). At large values  
 365 of  $Pe$ , diffusion is highly localized near the heterogeneity, causing large differences from the  
 366 ambient temperature in a narrow region. Moving toward smaller values of  $Pe$ , diffusion  
 367 becomes more efficient, and heat flows into the heterogeneity from a broad region. In the  
 368 spherical geometry, the volume of this region increases with the cube of distance from the  
 369 blob; when a large volume of ambient mantle contributes heat to the blob, the ambient

370 temperature change associated with that contribution is small. These two tendencies  
 371 are reflected by the trends in Figure 8, following the black curves from right to left.  
 372 At  $Pe \approx 1$ , where advective and diffusive heat transport are roughly in balance for the  
 373 spherical heterogeneity, we find a thermal perturbation that is relatively large in both  
 374 amplitude and extent. In Cartesian geometry, volume increases linearly with distance  
 375 from the tabular vein, and hence for a given inward heat flow, the thermal halo reaches  
 376 greater distances.

## 4.2. Dependence on dimensional size and upwelling rate

377 To conclude the discussion we reintroduce dimensional parameters and consider, inde-  
 378 pendently, the effect of changing the characteristic size  $R$  and the upwelling speed  $W$ .  
 379 Although these results could be deduced from earlier, dimensionless plots, they are pre-  
 380 sented in Figure 9 for clarity. To produce this figure, we have assumed values of material  
 381 parameters as given in Table 1. Panel a shows contours of degree of melting at  $t = 1$  for  
 382  $a = 1$  (analytical, black lines) and  $a = 1/4$  (numerical, red lines) for a spherical hetero-  
 383 geneity. Degree of melting is smallest at the top-right of the figure, for large blobs that  
 384 upwell rapidly and melt adiabatically, and largest at the bottom-left of the figure, for small  
 385 blobs that upwell slowly and melt in thermal equilibrium with their surroundings. Panel  
 386 b shows the dimensional temperature perturbation at a distance  $1.5 R$  from the centre of  
 387 the blob. For large blobs that upwell rapidly, loss of heat from the ambient mantle by  
 388 diffusion into the blob occurs in a region narrowly confined around the edge of the blob,  
 389 hence the temperature at  $r = 1.5 R$  is unaffected. For small blobs that upwell slowly, the  
 390 heat required to maintain thermal equilibrium with the ambient mantle is small, and it  
 391 is extracted over a very large volume around the blob, including  $r = 1.5 R$ . The case of



392 a tabular vein is shown in by contours of degree of melting and temperature perturba-  
 393 tion in panels c and d. The transition to diffusion-dominated melting is shifted to lower  
 394 Peclet numbers, meaning that tabular veins must be narrower or upwelling more slowly  
 395 to overcome their geometry and reach the same  $F_B$  as spherical blobs. Finally, note that  
 396 contours in all four panels have a slope of  $-2$ , which is consistent with the dependence of  
 397 Peclet number on  $R$  and  $W$  (note different  $x$ -scales in the two panels in Figure 9).

398 Results presented here are qualitatively consistent with those obtained by Sleep [1984].  
 399 Where his scaling analysis predicted a transition between thermally isolated and ther-  
 400 mally equilibrated at a characteristic size of about five kilometers for an upwelling rate  
 401 of 3 cm/yr, Figure 9c suggests that for a tabular vein, the transition occurs at a slightly  
 402 smaller size of about one kilometer. This difference is insignificant given the model as-  
 403 sumptions. More substantial differences can be noted in the amount of melting and the  
 404 enhancement factors obtained by Sleep [1984] compared to our work. His model consid-  
 405 ered only the linearised melting relationship  $a = 1$ , and used an isobaric productivity of  
 406  $0.02 \text{ K}^{-1}$  corresponding to a temperature difference between the liquidus and solidus of  
 407  $\Delta T = 50 \text{ K}$ . This is about a factor of five smaller than the liquidus–solidus difference  
 408 obtained empirically by Pertermann and Hirschmann [2003a]; furthermore, near-solidus  
 409 productivity for G2 pyroxenite is about  $0.001 \text{ K}^{-1}$ , a factor of 20 smaller than the value  
 410 used by Sleep [1984]. This difference, as prescribed by equation (26), explains the very  
 411 large enhancement factors ( $3\text{--}7\times$  adiabatic) that he obtained. Despite this enhancement  
 412 to melting, the melt fractions reported by Sleep [1984] are small relative to those obtained  
 413 here. This is because he computed melting curves for only 5 to 12 km of mantle ascent,  
 414 which again is small relative to the value of  $\sim 50 \text{ km}$  between the melting onset-depth of

415 pyroxenite and peridotite, as estimated by Pertermann and Hirschmann [2003a] and used  
416 here.

## 5. Summary and implications

417 In this manuscript we have presented new theory for the melting of fertile mantle hetero-  
418 geneities, and shown that a properly formulated Peclet number measures the importance  
419 of diffusion-driven versus adiabatic melting. Our results also show that a simplified the-  
420 ory that considers a uniformly melting spherical blob, rather than one that captures the  
421 gradient in  $F$  and  $T$  with radius, accurately models the gross behavior of the system, de-  
422 viating only for Peclet numbers near unity. For intermediate values of upwelling rate, our  
423 model predicts that uniformly melting spherical blobs  $\lesssim 5$  km in radius have substantial  
424 melt enhancement by diffusion; for tabular veins, this transition occurs at a smaller size  
425 of  $\lesssim 1$  km. These conclusions confirm and extend the scaling analysis of Sleep [1984].

426 Under the assumptions and parameter choices listed above, we have demonstrated that  
427 diffusion of heat into an upwelling, fertile heterogeneity can lead to an increase in the  
428 degree of partial melting by a factor of two, generating extents of melting of 50–80%  
429 beneath the bottom of the ambient melting region. Furthermore, we have shown that  
430 the thermal anomaly imprinted on the ambient mantle is confined to within about two  
431 times the characteristic size of the heterogeneity, and ranges down to about -60 K. This  
432 temperature difference is small relative to the absolute temperature of the mantle, but sig-  
433 nificant when compared with the temperature drop due to decompression melting beneath  
434 a ridge. As such, melting of the ambient mantle will be suppressed in the neighborhood  
435 of a heterogeneity, with the onset of ambient melting occurring at shallower depths. The

436 expected change in the overall degree of ambient mantle melting due to this effect is on  
437 the order of 1–2% for a pyroxenite fraction of 5% [Phipps Morgan, 2001].

438 A detailed consideration of the geochemical implications of variable melting of hetero-  
439 geneities based on their size and shape is beyond the scope of the present paper. It might  
440 be argued, for example, that larger heterogeneities, which melt to lesser extents, preserve  
441 residual garnet to shallower depth, and hence may impart a greater garnet signature  
442 than smaller heterogeneities. Such arguments are based on the details of the partition  
443 coefficients, the mineral mode of recycled oceanic crust, and the rate and style of melt  
444 segregation from pyroxenitic heterogeneities [e.g. Prytulak and Elliott, 2009]. Moreover,  
445 most geochemical models of the contribution of magma from recycled oceanic crust indi-  
446 cate that to preserve a distinctive geochemical signature, such melts must ascend rapidly,  
447 in chemical isolation from the ambient mantle. This is thought to occur either by hy-  
448 drofracture and propagation of dikes, or by reactive flow and transport through high-flux  
449 dunite channels [e.g. Kelemen et al., 1995; Lundstrom et al., 2000; Spiegelman and Kele-  
450 men, 2003; Elliott and Spiegelman, 2003; Kogiso et al., 2004].

451 In this context, we emphasize that because the rate of reactive melting is proportional  
452 to the vertical magmatic flux, a local excess of melt supplied to the melting region from  
453 below can induce reactive channelization [Hewitt, 2010; Liang et al., 2010]. This predic-  
454 tion, considered in light of the large extents of melting for recycled crust below the base  
455 of ambient-mantle melting regime, supports the hypothesis by Lundstrom et al. [2000]  
456 that melt released from a fertile heterogeneity could induce chemically isolated, channel-  
457 ized melt transport. Furthermore, it is possible that the pyroxenite-derived magmatic  
458 flux through a channel would reduce the local solidus temperature, cool the channel by

459 consumption of latent heat, and give rise to a diffusive flux of heat into the channel (as we  
 460 predict to occur around the heterogeneity). A cool diffusion-halo around a dunite channel  
 461 will suppress adjacent melting of mantle peridotite, further isolating the magma as it is  
 462 transported.

## Appendix A: Analytical solutions

### A1. Homogeneous melting of a spherical blob

When the degree of melting is linearly dependent on the homologous temperature ( $a = 1$  in (11)), analytical methods can be used to obtain solutions. For the homogeneous blob, the problem is to solve for  $\theta(r, t)$  satisfying (18),

$$\frac{\partial \theta}{\partial t} = \frac{1}{\text{Pe}} \frac{1}{r^2} \frac{\partial}{\partial r} \left( r^2 \frac{\partial \theta}{\partial r} \right), \quad (\text{A1})$$

with boundary conditions

$$-\frac{\mathcal{S}}{\Delta\theta} + \left( 1 + \frac{\mathcal{S}}{\Delta\theta} \right) \frac{\partial \theta}{\partial t}(1, t) = \frac{3}{\text{Pe}} \frac{\partial \theta}{\partial r}(1, t), \quad (\text{A2})$$

$$\theta(\infty, t) = 0, \quad (\text{A3})$$

and initial condition

$$\theta(r, 0) = 0. \quad (\text{A4})$$

The first of the two boundary conditions represents the heat balance between the heat flowing into the blob due to the temperature gradient outside, the heat used to melt the blob (latent heat), and the heat used to raise the temperature of the blob (sensible heat). The expression (A2) follows directly from (10), (11) and (12) when  $a = 1$  (and also follows from partial differentiation with respect to  $t$  of the integral boundary condition described by (17) and (19a)). The second boundary condition states that the far-field temperature is constant (i.e. neglecting adiabatic decompression effects: see A5 for discussion of these

effects). The total degree of melting of the blob can be obtained from (9a), (10) and (11)

as

$$F_B(t) = \frac{\theta(1, t) + t}{\Delta\theta}. \quad (\text{A5})$$

In order to simplify later algebra, it is helpful to rescale time and temperature by the Peclet number and introduce a parameter  $\lambda$  as

$$\theta' = \frac{\theta}{\text{Pe}}, \quad t' = \frac{t}{\text{Pe}}, \quad \lambda = \frac{\Delta\theta}{\mathcal{S} + \Delta\theta}. \quad (\text{A6})$$

With this new scaling, the problem becomes

$$\frac{\partial\theta'}{\partial t'} = \frac{1}{r^2} \frac{\partial}{\partial r} \left( r^2 \frac{\partial\theta'}{\partial r} \right), \quad (\text{A7})$$

with boundary conditions and initial condition

$$\frac{\partial\theta'}{\partial t'}(1, t') = -1 + \lambda + 3\lambda \frac{\partial\theta'}{\partial r}(1, t'), \quad (\text{A8})$$

$$\theta'(\infty, t') = 0, \quad (\text{A9})$$

$$\theta'(r, 0) = 0, \quad (\text{A10})$$

and

$$F_B(t') = \frac{\text{Pe}}{\Delta\theta} (\theta'(1, t') + t'). \quad (\text{A11})$$

<sup>463</sup> For the remainder of this appendix, we will drop the primes and use the rescaled variables.

**Laplace transform solution:** The governing partial differential equation (A7) can be simplified by introducing a new variable  $u(r, t)$  as

$$\theta(r, t) = \frac{u(r, t)}{r}, \quad (\text{A12})$$

to give

$$\frac{\partial u}{\partial t} = \frac{\partial^2 u}{\partial r^2} \quad (\text{A13})$$

with boundary conditions and initial condition

$$\frac{\partial u}{\partial t}(1, t) = -1 + \lambda + 3\lambda \left( \frac{\partial u}{\partial r}(1, t) - u(1, t) \right), \quad (\text{A14})$$

$$u(r, t)/r \rightarrow 0 \text{ as } r \rightarrow \infty, \quad (\text{A15})$$

$$u(r, 0) = 0. \quad (\text{A16})$$

Introduce the Laplace transform in time as

$$\tilde{u}(r, s) = \int_0^\infty u(r, t)e^{-st} dt. \quad (\text{A17})$$

The transformed problem is then

$$s\tilde{u} = \frac{\partial^2 \tilde{u}}{\partial r^2}, \quad (\text{A18})$$

with boundary conditions

$$s\tilde{u}(1, s) = \frac{-1 + \lambda}{s} + 3\lambda \left( \frac{\partial \tilde{u}}{\partial r}(1, s) - \tilde{u}(1, s) \right), \quad (\text{A19})$$

$$\tilde{u}(r, s)/r \rightarrow 0 \text{ as } r \rightarrow \infty. \quad (\text{A20})$$

(A18) and (A20) imply

$$\tilde{u}(r, s) = A(s)e^{-\sqrt{s}(r-1)}, \quad (\text{A21})$$

for some function  $A(s)$  to be determined. (A19) then becomes

$$sA(s) = \frac{-1 + \lambda}{s} + 3\lambda (-\sqrt{s}A(s) - A(s)) \quad (\text{A22})$$

which gives  $A(s)$  as

$$A(s) = \frac{-1 + \lambda}{s(s + 3\lambda\sqrt{s} + 3\lambda)}. \quad (\text{A23})$$

Hence we have

$$\tilde{\theta}(r, s) = \frac{(\lambda - 1)e^{-\sqrt{s}(r-1)}}{rs(s + 3\lambda\sqrt{s} + 3\lambda)}. \quad (\text{A24})$$

To find  $\theta(r, t)$  we need to find the inverse Laplace transform of the above function. This can be obtained by factoring the denominator, splitting into partial fractions, and performing the inverse Laplace transform term-by-term. Writing

$$\frac{1}{s(s + 3\lambda\sqrt{s} + 3\lambda)} \equiv \frac{1}{s(\sqrt{s} + a_\lambda)(\sqrt{s} + b_\lambda)} \quad (\text{A25})$$

where

$$a_\lambda = \frac{3\lambda + \sqrt{9\lambda^2 - 12\lambda}}{2}, \quad b_\lambda = \frac{3\lambda - \sqrt{9\lambda^2 - 12\lambda}}{2}, \quad (\text{A26})$$

the inverse Laplace transform of  $\tilde{\theta}(r, s)$  is

$$\theta(r, t) = \frac{\lambda - 1}{r} \left[ \frac{1}{a_\lambda b_\lambda} \operatorname{erfc}(\eta) + \frac{e^{-\eta^2}}{a_\lambda - b_\lambda} \left( \frac{1}{a_\lambda} w(i\eta + ia_\lambda\sqrt{t}) - \frac{1}{b_\lambda} w(i\eta + ib_\lambda\sqrt{t}) \right) \right], \quad (\text{A27})$$

where  $\eta = (r - 1)/(2\sqrt{t})$  and  $w(z) = e^{-z^2} \operatorname{erfc}(-iz)$  is the Faddeeva function. (A27) is the analytical solution to the linear homogeneous blob problem, and can be calculated rapidly with the aid of efficient routines for calculating the Faddeeva function [Weideman, 1994]. A solution similar to the above was recently obtained by Oliver [2008] for a related problem of spherical heat generation and conduction. The corresponding degree of melting is given by (A11),

$$F_B(t) = \frac{\text{Pe}}{\Delta\theta} \left( t + (\lambda - 1) \left[ \frac{1}{a_\lambda b_\lambda} + \frac{1}{a_\lambda - b_\lambda} \left( \frac{1}{a_\lambda} w(ia_\lambda\sqrt{t}) - \frac{1}{b_\lambda} w(ib_\lambda\sqrt{t}) \right) \right] \right), \quad (\text{A28})$$

and is plotted in Figure 3.

The expressions in (A27) and (A28) are somewhat cumbersome to work with when studying the asymptotic behaviors of the solution. It is easier to study the asymptotic behavior of the Laplace transform solution in  $s$ , and then relate the asymptotics in  $s$  to the asymptotics in  $t$  (see A5). For this purpose, note that  $\tilde{F}_B(s)$  is given by the expression

$$\tilde{F}_B(s) = \frac{\text{Pe}}{\Delta\theta} \left( \frac{1}{s^2} + \tilde{\theta}(1, s) \right) = \frac{\text{Pe}}{\Delta\theta} \left( \frac{1}{s^2} + A(s) \right). \quad (\text{A29})$$

## A2. Radially variable melting of a spherical blob

If we do not assume that the blob is homogeneous, and instead allow it to have a radial temperature profile due to the conduction of heat through the blob, then we must solve the heat conservation equation both inside and outside the blob. Using the rescaled variables of (A6), the governing equations are

$$\frac{\partial \theta}{\partial t} = -1 + \lambda + \frac{\lambda}{r^2} \frac{\partial}{\partial r} \left( r^2 \frac{\partial \theta}{\partial r} \right), \quad 0 \leq r < 1, \quad (\text{A30})$$

$$\frac{\partial \theta}{\partial t} = \frac{1}{r^2} \frac{\partial}{\partial r} \left( r^2 \frac{\partial \theta}{\partial r} \right), \quad r > 1. \quad (\text{A31})$$

On the surface of the blob, both the temperature and the heat flux must be continuous, i.e.

$$\theta(r, t), \frac{\partial \theta}{\partial r}(r, t) \text{ continuous on } r = 1, \quad (\text{A32})$$

and as before the initial condition is

$$\theta(r, 0) = 0, \quad (\text{A33})$$

and the boundary condition in the far-field is

$$\theta(\infty, t) = 0. \quad (\text{A34})$$

**Laplace transform solution:** As before, the governing partial differential equations can be simplified by writing

$$\theta(r, t) = \frac{u(r, t)}{r} \quad (\text{A35})$$

to obtain

$$\frac{\partial u}{\partial t} = (\lambda - 1)r + \lambda \frac{\partial^2 u}{\partial r^2}, \quad 0 \leq r < 1, \quad (\text{A36})$$

$$\frac{\partial u}{\partial t} = \frac{\partial^2 u}{\partial r^2}, \quad r > 1, \quad (\text{A37})$$



with boundary and initial conditions

$$u(r, t), \frac{\partial u}{\partial r}(r, t) \quad \text{continuous on } r = 1, \quad (\text{A38})$$

$$u(r, t)/r \quad \text{finite as } r \rightarrow 0, \quad (\text{A39})$$

$$u(r, t)/r \rightarrow 0 \quad \text{as } r \rightarrow \infty, \quad (\text{A40})$$

$$u(r, 0) = 0. \quad (\text{A41})$$

The Laplace transformed problem is

$$s\tilde{u} = \frac{(\lambda - 1)r}{s} + \lambda \frac{\partial^2 \tilde{u}}{\partial r^2}, \quad 0 \leq r < 1, \quad (\text{A42})$$

$$s\tilde{u} = \frac{\partial^2 \tilde{u}}{\partial r^2}, \quad r > 1, \quad (\text{A43})$$

with boundary conditions

$$\tilde{u}(r, s), \frac{\partial \tilde{u}}{\partial r}(r, s) \quad \text{continuous on } r = 1, \quad (\text{A44})$$

$$\tilde{u}(r, s)/r \quad \text{finite as } r \rightarrow 0, \quad (\text{A45})$$

$$\tilde{u}(r, s)/r \rightarrow 0 \quad \text{as } r \rightarrow \infty. \quad (\text{A46})$$

The governing equations (A42) and (A43) can be integrated using the boundary conditions (A45) and (A46) to give

$$\tilde{u}(r, s) = \begin{cases} \frac{(\lambda - 1)r}{s^2} + B(s) \sinh\left(\sqrt{\frac{s}{\lambda}}r\right), & 0 \leq r < 1, \\ C(s)e^{-\sqrt{s}(r-1)}, & r > 1. \end{cases} \quad (\text{A47})$$

The two functions  $B(s)$  and  $C(s)$  are determined by the continuity requirements of (A44)

$$B(s) = -\frac{\lambda - 1}{s^2} \left( \frac{1 + \sqrt{s}}{\sqrt{\frac{s}{\lambda}} \cosh \sqrt{\frac{s}{\lambda}} + \sqrt{s} \sinh \sqrt{\frac{s}{\lambda}}} \right), \quad (\text{A48})$$

$$C(s) = \frac{\lambda - 1}{s^2} \left( \frac{\sqrt{\frac{s}{\lambda}} \cosh \sqrt{\frac{s}{\lambda}} - \sinh \sqrt{\frac{s}{\lambda}}}{\sqrt{\frac{s}{\lambda}} \cosh \sqrt{\frac{s}{\lambda}} + \sqrt{s} \sinh \sqrt{\frac{s}{\lambda}}} \right). \quad (\text{A49})$$

The solution for  $\tilde{\theta}(r, s)$  is thus

$$\tilde{\theta}(r, s) = \begin{cases} \frac{(\lambda - 1)}{s^2} + \frac{B(s)}{r} \sinh\left(\sqrt{\frac{s}{\lambda}}r\right), & 0 \leq r < 1, \\ \frac{C(s)}{r} e^{-\sqrt{s}(r-1)}, & r > 1. \end{cases} \quad (\text{A50})$$

465 To find  $\theta(r, t)$  we must obtain the inverse Laplace transform of the above function. Unfor-  
 466 tunately, there does not appear to be a simple analytical inverse of (A50). However, the  
 467 inverse can be calculated numerically using efficient routines for numerical inverse Laplace  
 468 transforms [de Hoog et al., 1982; Hollenbeck, 1998].

The degree of melting within the blob is given by

$$F(r, t) = \frac{\text{Pe}}{\Delta\theta} (t + \theta(r, t)), \quad (\text{A51})$$

and hence using (A50) we have

$$\tilde{F}(r, s) = \frac{\text{Pe}}{\Delta\theta} \left( \frac{1}{s^2} + \tilde{\theta}(r, s) \right) = \frac{\text{Pe}}{\Delta\theta} \left( \frac{\lambda}{s^2} + \frac{B(s)}{r} \sinh\sqrt{\frac{s}{\lambda}}r \right). \quad (\text{A52})$$

The Laplace transform of the mean degree of melting  $\bar{F}(t)$  is thus

$$\begin{aligned} \tilde{\bar{F}}(s) &= 3 \int_0^1 \tilde{F}(r, s) r^2 dr \\ &= \frac{\text{Pe}}{\Delta\theta} \left( \frac{\lambda}{s^2} + \frac{3B(s)\lambda}{s} \left( \sqrt{\frac{s}{\lambda}} \cosh\sqrt{\frac{s}{\lambda}} - \sinh\sqrt{\frac{s}{\lambda}} \right) \right) \\ &= \frac{\text{Pe}\lambda}{\Delta\theta} \left( \frac{1}{s^2} - \frac{3}{s}(1 + \sqrt{s})C(s) \right). \end{aligned} \quad (\text{A53})$$

469  $\bar{F}(t)$  can be obtained by finding the inverse Laplace transform of the above function.  
 470 This was done numerically using the routines of Hollenbeck [1998] to produce the profiles  
 471 plotted in Figure 5.

### A3. Melting of a tabular vein

In this section we briefly derive the tabular equivalents of the solutions given in A1 and A2. The general method of solution for a tabular geometry is identical to that for a

spherical geometry. The only change that needs to be made is that the Laplacian operator is now given by

$$\nabla^2 = \frac{\partial^2}{\partial x^2} \quad (\text{A54})$$

where  $0 < x < 1$  is inside the sheet, and  $x > 1$  is outside to sheet. Symmetry is assumed about the origin, so that

$$\frac{\partial \theta}{\partial x} = 0 \quad \text{on } x = 0. \quad (\text{A55})$$

### 472 **A3.1. Homogeneous melting**

The problem is

$$\frac{\partial \theta}{\partial t} = \frac{\partial^2 \theta}{\partial x^2} \quad (\text{A56})$$

with boundary conditions and initial condition

$$\frac{\partial \theta}{\partial t}(1, t) = -1 + \lambda + \lambda \frac{\partial \theta}{\partial x}(1, t), \quad (\text{A57})$$

$$\theta(\infty, t) = 0, \quad (\text{A58})$$

$$\theta(x, 0) = 0. \quad (\text{A59})$$

The Laplace transformed problem is

$$s\tilde{\theta} = \frac{\partial^2 \tilde{\theta}}{\partial x^2}, \quad (\text{A60})$$

with boundary conditions

$$s\tilde{\theta}(1, s) = \frac{-1 + \lambda}{s} + \lambda \frac{\partial \tilde{\theta}}{\partial x}(1, s), \quad (\text{A61})$$

$$\tilde{\theta}(\infty, t) = 0, \quad (\text{A62})$$

and solution

$$\tilde{\theta}(x, s) = \frac{(\lambda - 1) e^{-\sqrt{s}(x-1)}}{s^{3/2} (\sqrt{s} + \lambda)}. \quad (\text{A63})$$

The inverse Laplace transform of this is

$$\theta(x, t) = \frac{\lambda - 1}{\lambda^2} \left[ e^{-\eta^2} \left( w \left( i\eta + i\lambda\sqrt{t} \right) + \frac{2\lambda\sqrt{t}}{\sqrt{\pi}} \right) - \left( 1 + 2\lambda\sqrt{t}\eta \right) \operatorname{erfc}(\eta) \right], \quad (\text{A64})$$

where  $\eta = (x - 1)/(2\sqrt{t})$  and  $w(z)$  is the Faddeeva function. The above solution can also be found in Carslaw and Jaeger [1959] (their equation (12) in section 12.4). The corresponding degree of melting is given by

$$F_B(t) = \frac{\text{Pe}}{\Delta\theta} \left( t + \frac{\lambda - 1}{\lambda^2} \left[ w \left( i\lambda\sqrt{t} \right) + \frac{2\lambda\sqrt{t}}{\sqrt{\pi}} - 1 \right] \right). \quad (\text{A65})$$

### 473 A3.2. Laterally variable melting

The governing equations are

$$\frac{\partial\theta}{\partial t} = -1 + \lambda + \frac{\partial^2\theta}{\partial x^2}, \quad 0 \leq x < 1, \quad (\text{A66})$$

$$\frac{\partial\theta}{\partial t} = \frac{\partial^2\theta}{\partial x^2}, \quad x > 1. \quad (\text{A67})$$

The Laplace transformed problem is

$$s\tilde{\theta} = \frac{\lambda - 1}{s} + \lambda \frac{\partial^2\tilde{\theta}}{\partial x^2}, \quad 0 \leq x < 1, \quad (\text{A68})$$

$$s\tilde{\theta} = \frac{\partial^2\tilde{\theta}}{\partial x^2}, \quad x > 1. \quad (\text{A69})$$

with solution

$$\tilde{\theta}(x, s) = \begin{cases} \frac{\lambda - 1}{s^2} + B(s) \cosh \left( \sqrt{\frac{s}{\lambda}} x \right), & 0 \leq x < 1, \\ C(s) e^{-\sqrt{s}(x-1)}, & x > 1. \end{cases} \quad (\text{A70})$$

The two functions  $B(s)$  and  $C(s)$  are determined by continuity as

$$B(s) = -\frac{\lambda - 1}{s^2} \left( \frac{1}{\cosh \sqrt{\frac{s}{\lambda}} + \frac{1}{\sqrt{\lambda}} \sinh \sqrt{\frac{s}{\lambda}}} \right), \quad (\text{A71})$$

$$C(s) = \frac{\lambda - 1}{s^2} \left( \frac{\frac{1}{\sqrt{\lambda}} \sinh \sqrt{\frac{s}{\lambda}}}{\cosh \sqrt{\frac{s}{\lambda}} + \frac{1}{\sqrt{\lambda}} \sinh \sqrt{\frac{s}{\lambda}}} \right). \quad (\text{A72})$$

The Laplace transform of the mean degree of melting  $\bar{F}(t)$  is

$$\begin{aligned}\tilde{\bar{F}}(s) &= \int_0^1 \tilde{F}(x, s) dx \\ &= \frac{\text{Pe}}{\Delta\theta} \left( \frac{\lambda}{s^2} + B(s) \sqrt{\frac{\lambda}{s}} \sinh \sqrt{\frac{s}{\lambda}} \right) \\ &= \frac{\text{Pe}\lambda}{\Delta\theta} \left( \frac{1}{s^2} - \frac{C(s)}{\sqrt{s}} \right).\end{aligned}\tag{A73}$$

#### A4. Leading order asymptotics

474 All the problems considered in this manuscript (both spherical/tabular and homoge-  
475 neous/radially varying) have the same leading order behavior for large and small  $t$ . This  
476 behavior is exactly that which is expected from a simple thermodynamic analysis of the  
477 two extremes of a thermally isolated blob and a blob in thermal equilibrium with the  
478 ambient mantle. This leading order behavior has been described by Sleep [1984] and is  
479 depicted in Figure 2.

The large- $t$  and small- $t$  asymptotic behavior of  $F_B(t)$  and  $\bar{F}(t)$  can be obtained directly from the asymptotic behavior of the Laplace transforms  $\tilde{F}_B(s)$  and  $\tilde{\bar{F}}(s)$  for small  $s$  and large  $s$  respectively. The leading order asymptotics of  $\tilde{F}_B(s)$  are given by series expansion of (A29) as

$$\tilde{F}_B(s) \sim \begin{cases} \frac{\text{Pe}\lambda}{\Delta\theta s^2} + \mathcal{O}\left(\frac{1}{s^{5/2}}\right), & \text{for } s \gg 1, \\ \frac{\text{Pe}}{\Delta\theta s^2} + \mathcal{O}\left(\frac{1}{s}\right), & \text{for } s \ll 1.\end{cases}\tag{A74}$$

The leading order asymptotics of  $\tilde{\bar{F}}(s)$  from (A53) are identical. By inverse Laplace transforming term-by-term we obtain the leading order asymptotics of  $F_B(t)$  as

$$F_B(t) \sim \begin{cases} \frac{\text{Pe}\lambda t}{\Delta\theta} + \mathcal{O}(t^{3/2}), & \text{for } t \ll 1, \\ \frac{\text{Pe}t}{\Delta\theta} + \mathcal{O}(1), & \text{for } t \gg 1.\end{cases}\tag{A75}$$

The above expression can be written in dimensional units using (2) as

$$F_B(t) \sim \begin{cases} \frac{p_0 - p(t)}{\gamma(\Delta T + L/c_p)} + \mathcal{O}(t^{3/2}), & \text{for } t \ll R^2/\kappa, \\ \frac{p_0 - p(t)}{\gamma\Delta T} + \mathcal{O}(1), & \text{for } t \gg R^2/\kappa, \end{cases} \quad (\text{A76})$$

480 which agrees with the simple thermodynamic analysis of Sleep [1984] (his equations 9  
481 and 10 respectively, see Figure 2). Higher order asymptotic expansions for  $F_B(t)$  can be  
482 obtained by considering higher order terms in the series expansions of (A74). Differences  
483 between  $F_B(t)$  and  $\bar{F}(t)$ , and the tabular and spherical geometries, become evident with  
484 the inclusion of higher order terms.

### A5. Adiabatic decompression effects

Up to this point the effects of adiabatic decompression have been ignored, as it has been assumed that the far-field temperature of the ambient mantle is constant. In fact, the far-field temperature of the ambient mantle will decrease as the pressure decreases as a consequence of adiabatic decompression. It is straightforward to include this effect, at least in a linearized sense. If the temperature differences are small, such that  $(T_0 - T_1)/T_0 \ll 1$ , the adiabatic decompression term in the energy equation (5) can be approximated as

$$-\alpha T \frac{\partial p}{\partial t} \approx -\alpha T_0 \frac{\partial p}{\partial t}. \quad (\text{A77})$$

The non-dimensional radially varying blob problem is then

$$\frac{\partial \theta}{\partial t} = -\lambda \mathcal{A} - 1 + \lambda + \frac{\lambda}{r^2} \frac{\partial}{\partial r} \left( r^2 \frac{\partial \theta}{\partial r} \right), \quad 0 \leq r < 1, \quad (\text{A78})$$

$$\frac{\partial \theta}{\partial t} = -\mathcal{A} + \frac{1}{r^2} \frac{\partial}{\partial r} \left( r^2 \frac{\partial \theta}{\partial r} \right), \quad r > 1, \quad (\text{A79})$$

where  $\mathcal{A}$  is the adiabatic parameter, defined by

$$\mathcal{A} = \frac{\alpha T_0 \gamma}{\rho c_p}. \quad (\text{A80})$$

The boundary conditions on the surface of the blob, and the initial condition are as before.

The far-field boundary condition becomes

$$\theta(\infty, t) = -\mathcal{A}t, \quad (\text{A81})$$

reflecting the fact that the far-field temperature drops as the blob ascends. By writing

$$\theta(r, t) = -\mathcal{A}t + (1 - \mathcal{A})\vartheta(r, t), \quad (\text{A82})$$

we recover the problem that has already been solved neglecting adiabatic decompression,

i.e.

$$\frac{\partial \vartheta}{\partial t} = -1 + \lambda + \frac{\lambda}{r^2} \frac{\partial}{\partial r} \left( r^2 \frac{\partial \vartheta}{\partial r} \right), \quad 0 \leq r < 1, \quad (\text{A83})$$

$$\frac{\partial \vartheta}{\partial t} = \frac{1}{r^2} \frac{\partial}{\partial r} \left( r^2 \frac{\partial \vartheta}{\partial r} \right), \quad r > 1, \quad (\text{A84})$$

$$\vartheta(\infty, t) = 0. \quad (\text{A85})$$

Hence to calculate the solution for a problem which includes the adiabatic decompression term, we simply find the solution without the term, and then use (A82). This works for both the homogeneous and the radially varying blob problems, as well as for the tabular geometry. Since

$$F(r, t) = \frac{\text{Pe}}{\Delta\theta} (t + \theta(r, t)) = \frac{\text{Pe}}{\Delta\theta} (1 - \mathcal{A}) (t + \vartheta(r, t)), \quad (\text{A86})$$

the melt productivity decreases by a factor of  $(1 - \mathcal{A})$  when the adiabatic decompression term is included. For example, the values of the dimensional degree of melting at the two extremes changes from that given by (A76) to

$$F_B(t) \sim \begin{cases} \frac{p_0 - p(t)}{\Delta T + L/c_p} \left( \frac{1}{\gamma} - \frac{\alpha T_0}{\rho c_p} \right) + \mathcal{O}(t^{3/2}), & \text{for } t \ll R^2/\kappa, \\ \frac{p_0 - p(t)}{\Delta T} \left( \frac{1}{\gamma} - \frac{\alpha T_0}{\rho c_p} \right) + \mathcal{O}(1), & \text{for } t \gg R^2/\kappa. \end{cases} \quad (\text{A87})$$

485 The first of the above two cases can be recognized as following directly from the usual

486 expression for the productivity in a one-dimensional decompression melting (e.g. Equation 11

487 (3.14) of Asimow et al. [1997]). The second case is in agreement with expressions for the  
 488 productivity assuming complete thermal equilibration (e.g. equation (19) of Phipps Mor-  
 489 gan [2001]).

## Appendix B: Numerical solutions

490 The governing equations with  $a < 1$  are nonlinear and must be solved numerically.  
 491 To do so we use a semi-implicit, centred-difference discretization on a non-uniform grid.  
 492 We solve the resulting system of nonlinear algebraic equations with a Newton-Krylov  
 493 (GMRES) scheme and an explicit LU preconditioner; these are provided by the Portable,  
 494 Extensible Toolkit for Scientific Computation [version 3.1; Balay et al., 2011, 2010; Katz  
 495 et al., 2007]. Details of the discretization are given in this appendix for the case of a  
 496 spherical heterogeneity; corresponding equations for the tabular vein are obtained in a  
 497 similar manner. Simulation code is available by email request to the first author.

The discretization is semi-implicit in time,

$$\frac{\theta^{n+1} - \theta^n}{\Delta t} = \frac{1}{2\text{Pe}} \left[ \left( \frac{1}{r^2} \frac{\partial}{\partial r} \left( r^2 \frac{\partial \theta}{\partial r} \right) \right)^{n+1} + \left( \frac{1}{r^2} \frac{\partial}{\partial r} \left( r^2 \frac{\partial \theta}{\partial r} \right) \right)^n \right], \quad (\text{B1})$$

498 where  $\Delta t$  is the time-step, chosen such that  $t_n = n\Delta t$  for  $n \in [0, N_t - 1]$ . Superscripts in  
 499 the semi-discrete equation (B1) refer to the time-step number.

Spatial derivatives are discretized with a centered difference scheme,

$$\frac{1}{r^2} \frac{\partial}{\partial r} \left( r^2 \frac{\partial \theta}{\partial r} \right) \approx \frac{\left[ \left( \frac{r_{i+1} + r_i}{2} \right)^2 \left( \frac{\theta_{i+1} - \theta_i}{r_{i+1} - r_i} \right) \right] - \left[ \left( \frac{r_i + r_{i-1}}{2} \right)^2 \left( \frac{\theta_i - \theta_{i-1}}{r_i - r_{i-1}} \right) \right]}{\frac{1}{2}(r_{i+1} - r_{i-1}) r_i^2}. \quad (\text{B2})$$

Values of the radius are specified at a set of discrete points  $i \in [0, N_r - 1]$  using

$$r_i = 1 + (r_{\max} - 1) \left( \frac{i}{N_r - 1} \right)^\xi, \quad (\text{B3})$$



500 where  $\xi \geq 1$  is a power that determines the relative concentration of grid-points near the  
 501 blob. We have found that  $\xi = 2$  provides a good balance between accuracy and speed of  
 502 numerical convergence.

The boundary condition at  $r \rightarrow \infty$  is a straightforward Dirichlet condition, which we  
 apply at  $r = r_{\max} \gg 1$ . The boundary condition at  $r = 1$  is more difficult. For  $F_B \leq 1$ ,  
 the semi-implicit discretization of this condition is as follows

$$\theta_0^n = -n\Delta t + \Delta\theta \Theta_B^n, \quad (\text{B4})$$

where  $\Theta_B^n$  is the dimensionless homologous temperature of the blob at the present step,  
 obtained with equation (15), which depends on the unknown value of  $f_B$ . The current  
 Newton iterate  $\check{f}_B$  can be calculated using the current Newton iterate of the solution  
 vector  $\check{\theta}_i^n$  by discrete integration with the trapezoidal rule,

$$\check{f}_B^n = f_b^{n-1} + \frac{\Delta t}{\Delta\theta + a\mathcal{S}} \left( 1 + \frac{3}{\text{Pe}} \frac{\check{\theta}_1^n + \theta_1^{n-1} - \check{\theta}_0^n - \theta_0^{n-1}}{2(r_1 - r_0)} \right). \quad (\text{B5})$$

503 The discrete boundary conditions and diffusion equation are then recast as equations  
 504 for the elements of the point-wise Newton residual vector of the current iterate  $\check{\varrho}_i$ . We  
 505 provide an analytical Jacobian matrix  $J_{ij} = \partial\check{\varrho}_i/\partial\check{\theta}_j$ , and the Newton scheme is iterated  
 506 until the residual vector satisfies  $\|\check{\varrho}_i\|_2 < \text{tol}$ . We use a tolerance of  $10^{-10}$ .

Numerical solutions can be compared with the analytical solution for the linear case,  
 $a = 1$ . Percent error is computed as

$$e = \frac{\|\theta_{\text{exact}} - \theta_{\text{numerical}}\|_2}{\|\theta_{\text{exact}}\|_2} \times 100, \quad (\text{B6})$$

507 where the 2-norm is calculated over all combinations of  $r_i, t_n$  used in the numerical model-  
 508 run. We obtain perfect second order convergence with grid-spacing for grids up to  $N_r =$

509 3200 with  $\xi = 2$ ; we obtain little improvement in accuracy for  $N_t \gtrsim 800$ . For a grid with  
510  $N_r = 1600$  and  $N_t = 1000$  we find that  $e = 0.003\%$ .

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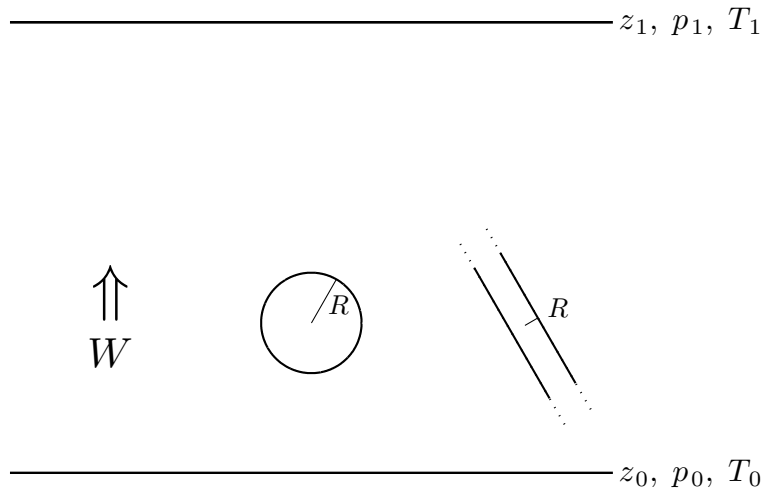
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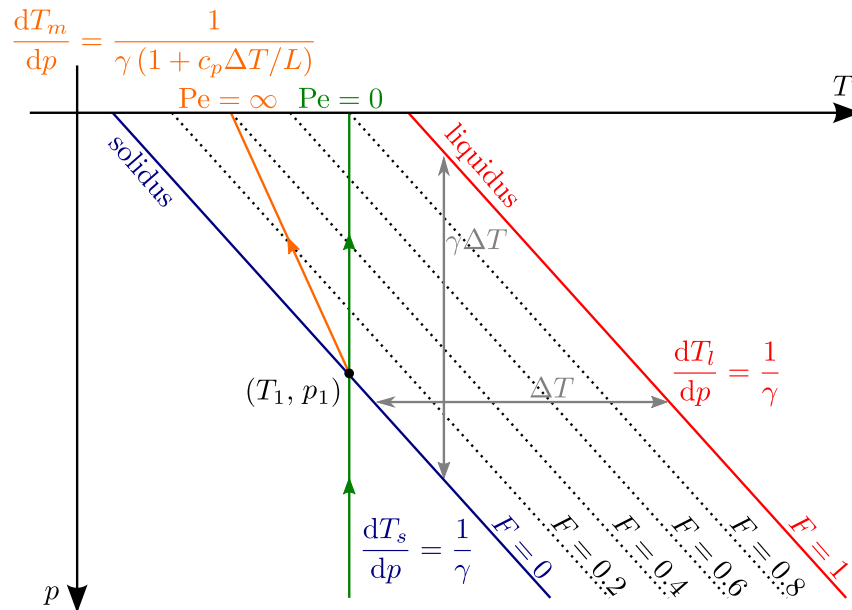
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**Figure 1.** Schematic diagram of the upwelling column with idealized fertile blob and fertile tabular vein. The blob is a sphere in three dimensions, while the tabular vein is an infinite sheet. The fertile material crosses its solidus at  $z = z_0$  and  $t = 0$ ; far-field ambient mantle crosses its solidus at  $z = z_1$  and  $t = t_1$ .

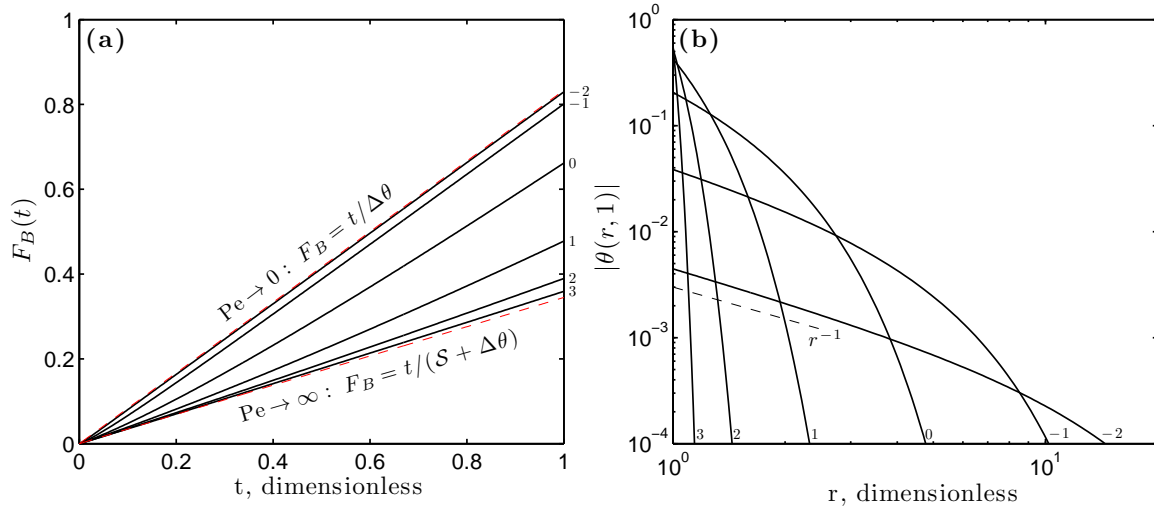




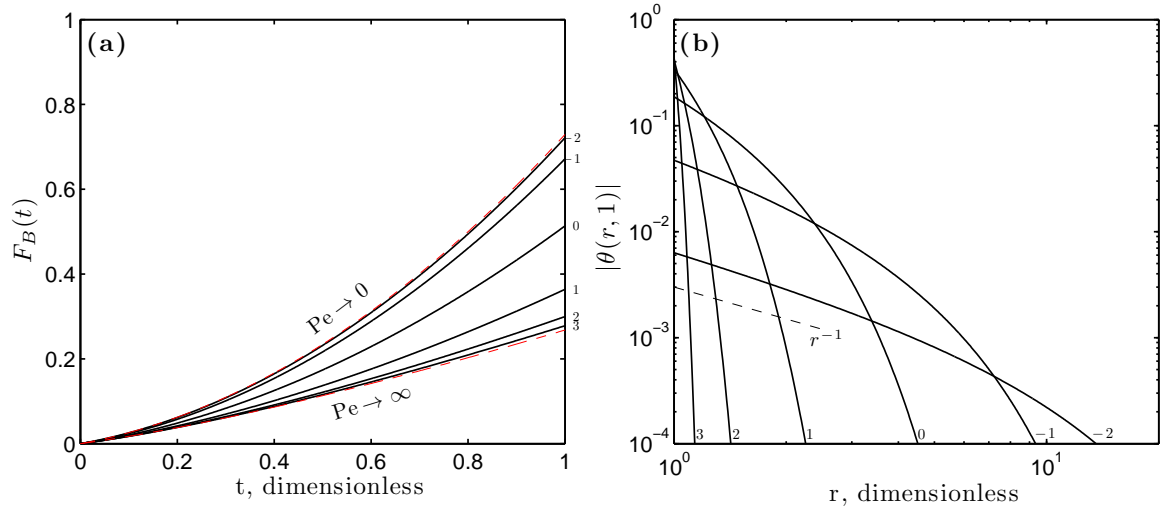
**Figure 2.** Thermodynamic analysis of the melting of the blob in the two extreme cases of complete thermal isolation ( $Pe = \infty$ ) and complete thermal equilibration ( $Pe = 0$ ). Diagram based on Sleep [1984], Fig. 5. Orange and green lines represent temperature-pressure paths for these extremes. Linear solidus and liquidus lines are assumed (red and blue lines), the dependence of melt fraction on homologous temperature is taken to be linear (i.e.  $a = 1$ ), and we neglect the adiabatic temperature gradient (i.e.  $\mathcal{A} = 0$ ). Degrees of melting are depicted by dotted contours.

Parameter	Value or range	Units	Comment
$a$	0.25 – 1	-	Coefficient in melting relation (11)
$c_p$	1200	J kg <sup>-1</sup> K <sup>-1</sup>	Specific heat capacity
$g$	10	m s <sup>-2</sup>	Gravitational acceleration
$k$	= $\rho c_p \kappa$	J K <sup>-1</sup> m <sup>-1</sup> s <sup>-1</sup>	Thermal conductivity
$L$	$4 \times 10^5$	J kg <sup>-1</sup>	Latent heat of melting
$R$	$10^{-2} - 10^3$	km	Characteristic size of heterogeneity
$T_0$	1623	K	Reference melting temperature
$W$	$10^{-1} - 10^3$	cm a <sup>-1</sup>	Upwelling rate
$\alpha$	$3 \times 10^{-5}$	K <sup>-1</sup>	Thermal expansivity
$\gamma$	$8.3 \times 10^6$	Pa K <sup>-1</sup>	Clausius-Clapeyron slope (120 K/GPa)
$\Delta p$	$1.7 \times 10^9$	Pa	Pressure interval (110–60 km depth)
$\Delta T$	250	K	$T$ -difference between liquidus and solidus
$\kappa$	$10^{-6}$	m <sup>2</sup> s <sup>-1</sup>	Thermal diffusivity
$\rho$	3300	kg m <sup>-3</sup>	Density
$\mathcal{A}$	0.1	-	Adiabatic parameter
$\Delta\theta$	1.2	-	Melting-temperature interval
Pe		-	Peclet number
$\mathcal{S}$	1.7	-	Stefan number

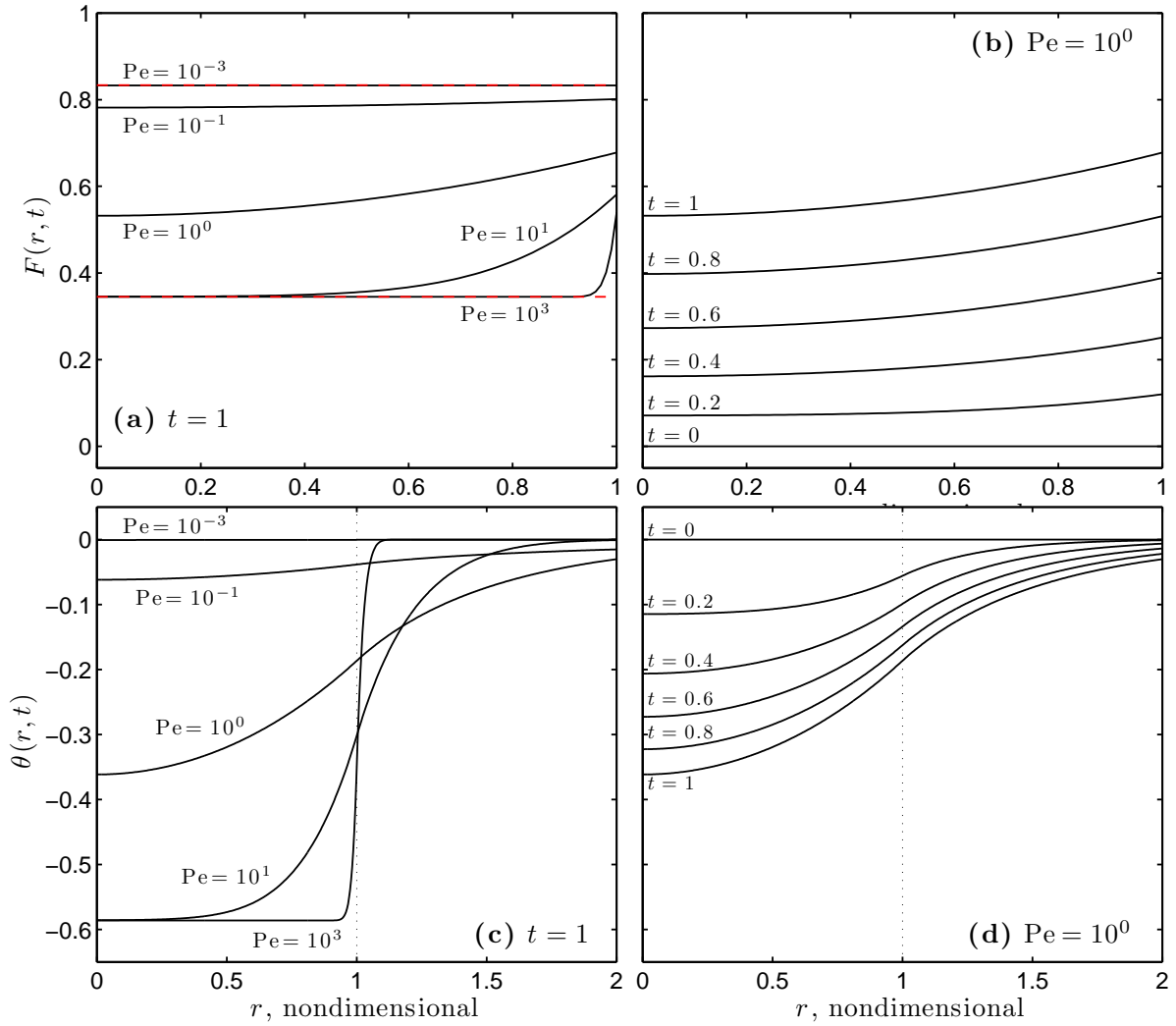
**Table 1.** Dimensional and non-dimensional parameters.



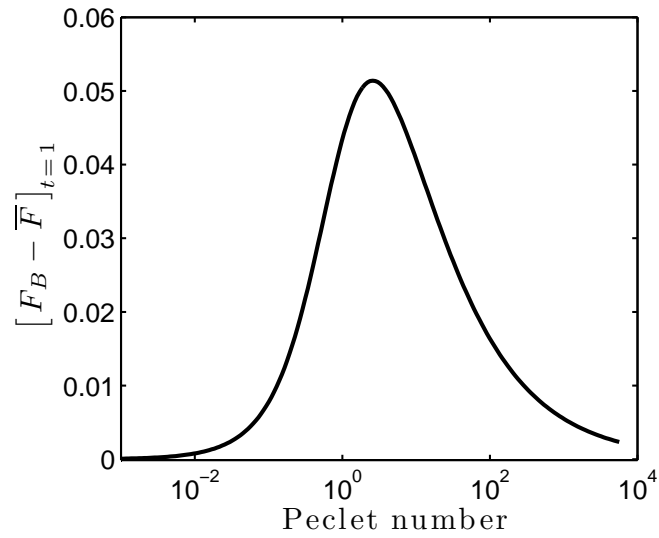
**Figure 3.** Analytical solution for equations (17–20) with  $a = 1$  and other parameters as in Table 1. Curves are labelled with  $\log_{10} \text{Pe}$ . **(a)** The mean degree of melting within the blob as a function of time, for different values of  $\text{Pe}$ . The red dashed curves are the limiting cases of  $\text{Pe} \rightarrow 0$  and  $\text{Pe} \rightarrow \infty$ . **(b)** The dimensionless temperature disturbance  $|\theta(r, 1)|$  caused by diffusion into the blob as a function of dimensionless radius at the dimensionless time  $t = 1$ .



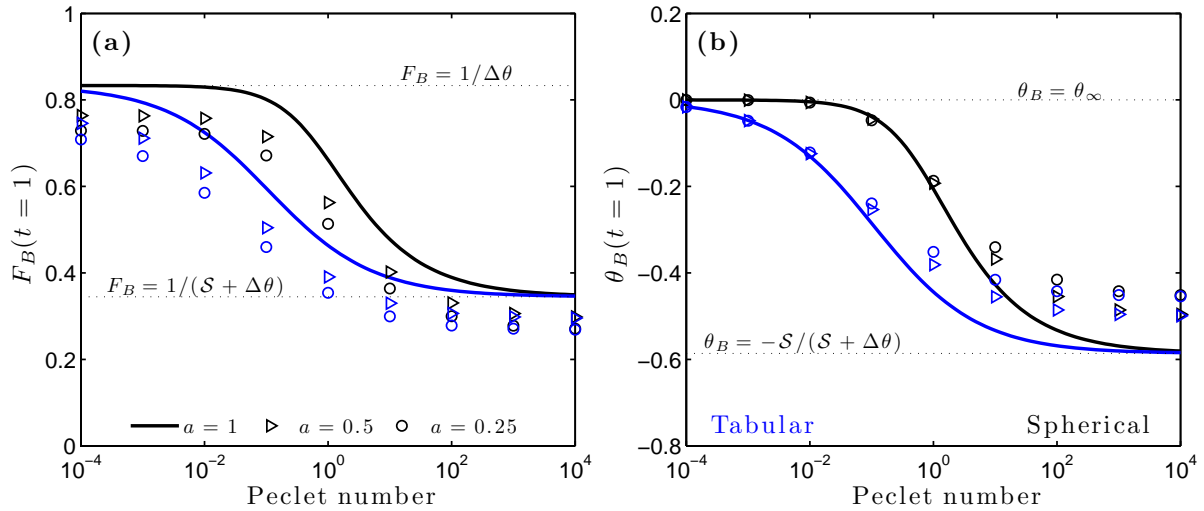
**Figure 4.** Numerical solutions for equations (15), (16), (18–20) with  $a = 1/4$  and other parameters as in Table 1. Each curve is labelled with  $\log_{10} Pe$ . Details of panels (a) and (b) as in Figure 3.



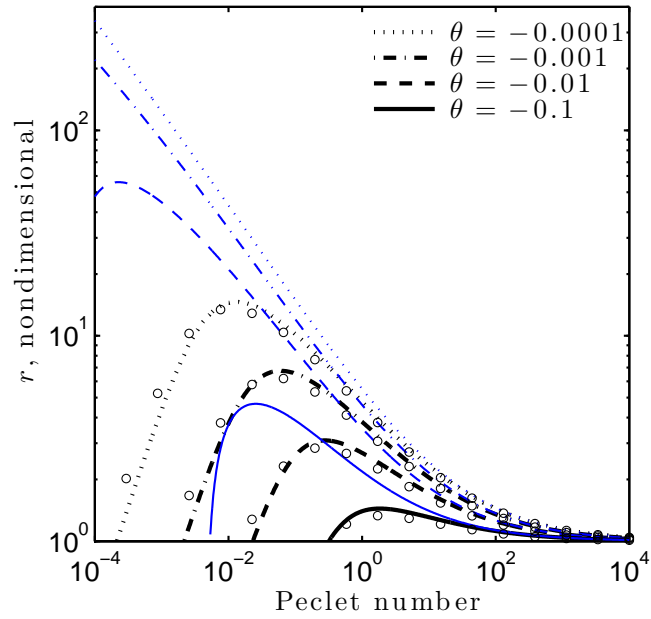
**Figure 5.** Radial profiles of melt fraction  $F(r, t)$  and temperature  $\theta(r, t)$  through the blob with  $a = 1$  (compare with Fig. 6 of Sleep [1984]). **(a)**  $F(r, t)$  at  $t = 1$  for different values of  $Pe$ . Red, dashed lines are the asymptotic values of  $F(r < 1, t)$  for  $Pe \rightarrow 0, \infty$ . **(b)**  $F(r, t)$  at  $Pe = 1$  for different values of  $t$ . **(c)**  $\theta(r, t)$  at  $t = 1$  for different values of  $Pe$ . **(d)**  $\theta(r, t)$  at  $Pe = 1$  for different values of  $t$ .



**Figure 6.** Difference between the radially averaged  $F_B(t)$  and the radially resolved  $\bar{F}(t)$  solutions for  $a = 1$  at time  $t = 1$ . For comparison, the radially resolved solution has been averaged over the blob. The temperature difference between the two models (not shown) is a factor of  $\Delta\theta$  larger than the difference in degree of melting.

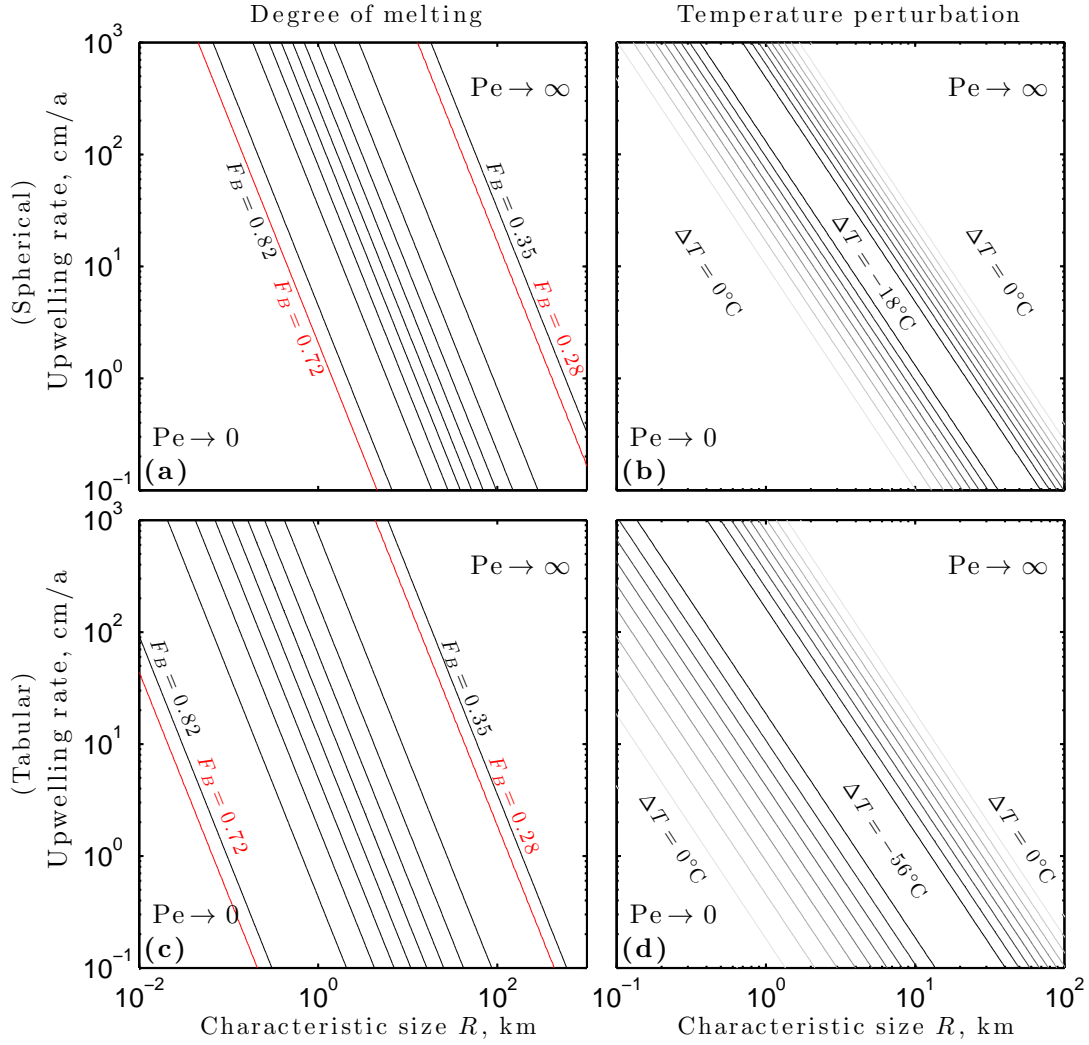


**Figure 7.** Summary of degree of melting (panel (a)) and dimensionless temperature (panel (b)) at  $t = 1$  for three values of  $a$ , computed for a range in Peclet number spanning the transition between asymptotic values. Black lines and symbols correspond to the spherical-blob model; blue lines and symbols correspond to the tabular-vein model. Dotted lines show asymptotic values as labelled. Note that in interpreting  $Pe$  for the tabular vein,  $R$  is the half-width, as shown in Figure 1.



**Figure 8.** The non-dimensional size of the diffusively cooled halo around the spherical blob (black) and the tabular vein (blue) at  $t = 1$ , as a function of the Peclet number. Lines are calculated with the analytical solution for  $a = 1$ ; each line represents the radius at which  $\theta(r, 1)$  reaches a specified value (see legend for values). Points are derived from spherical-blob simulations with  $a = 1/4$ ; their close correspondence with the lines indicates the associated value of  $\theta(r, 1)$ .





**Figure 9.** Contour plots of dimensional quantities as a function of characteristic heterogeneity size  $R$  in km and upwelling rate  $W$  in cm/yr. Other dimensional parameters as in Table 1. Panels (a) and (c) show degree of melting for the spherical blob and tabular vein, respectively. Black contours are computed with the analytical solution and  $a = 1$ ; red contours are computed with the numerical simulation and  $a = 1/4$ . Degree of melting is maximal in the bottom left corner and minimal in the top-right corner. Contour spacing is linear and the maximum (minimum) contour values are 1% less (more) than the appropriate asymptotic values. Asymptotic values for  $F_B$  are obtained using equations (24) and (25) with equation (11). For the nonlinear melting solution, only the maximum and minimum contours are shown. Panels (b) and (d) show the temperature perturbation at a distance  $1.5 R$  from the centre of the spherical blob or tabular vein, respectively. The largest absolute perturbation is  $\sim 20^\circ\text{C}$  (spherical) and  $\sim 56^\circ\text{C}$  (tabular) at this radius. Only the linear  $a = 1$  solution is plotted for temperature. As in previous figures, the results here exclude the background adiabatic gradient (see A5 for details).